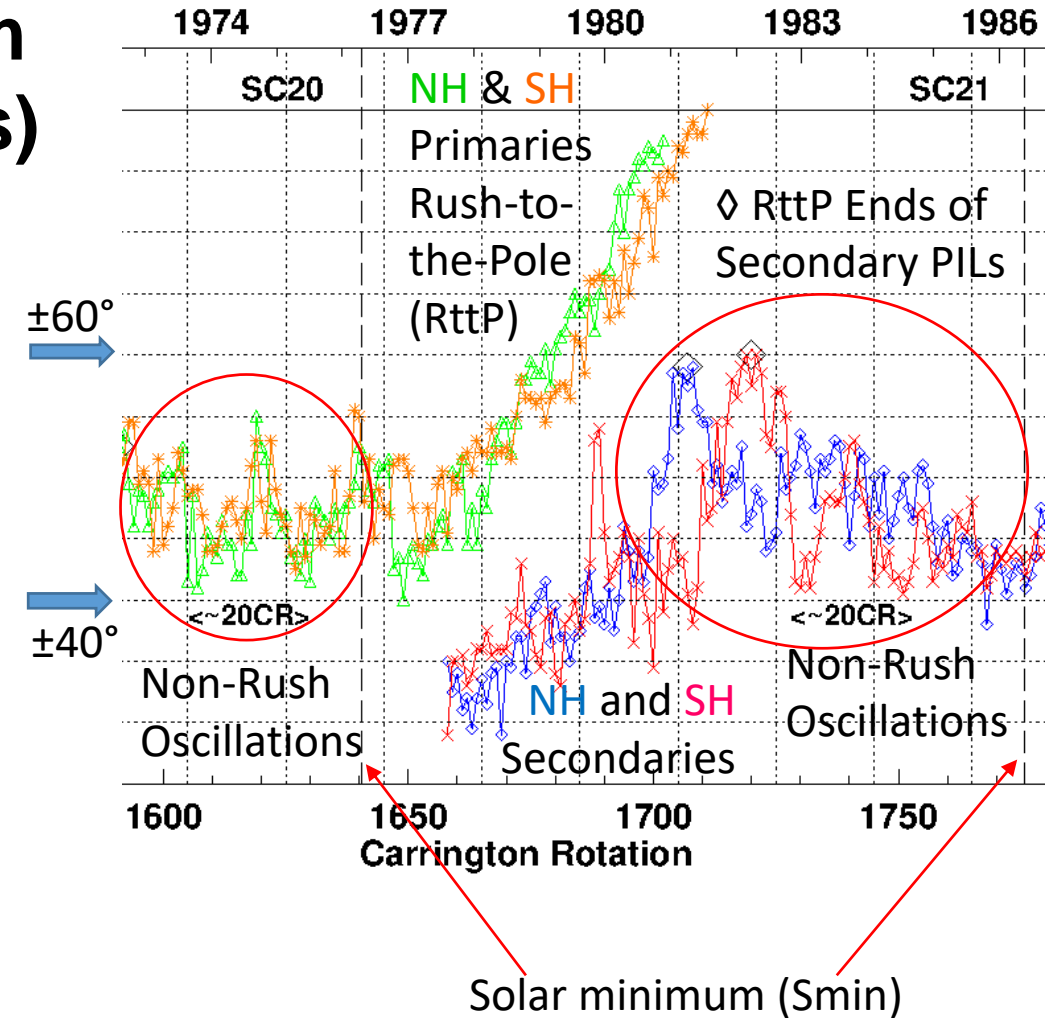


# SH006-06 Oscillations in Secondary to Primary Polar Crown Polarity Inversion Lines (PILs) around Solar Maximum (Smax) over Five Solar Cycles (SCs)

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 S. E. Gibson (1), I. M. Hewins (2),  
 R. H. McFadden (2), T. A. Kuchar (3)

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- (2) Institute for Scientific Research (ISR),  
 Boston College at HAO/NCAR
- (3) ISR, Boston College, Chestnut Hill, MA

NH=Northern Hemisphere  
 SH=Southern Hemisphere



# McIntosh Archive of 55 years of PILs (Dec'54-Aug'09, SC 19-23) and 35 years of Coronal Holes (CHs, Apr'74-Aug'09)

McIntosh Archive Synoptic Map

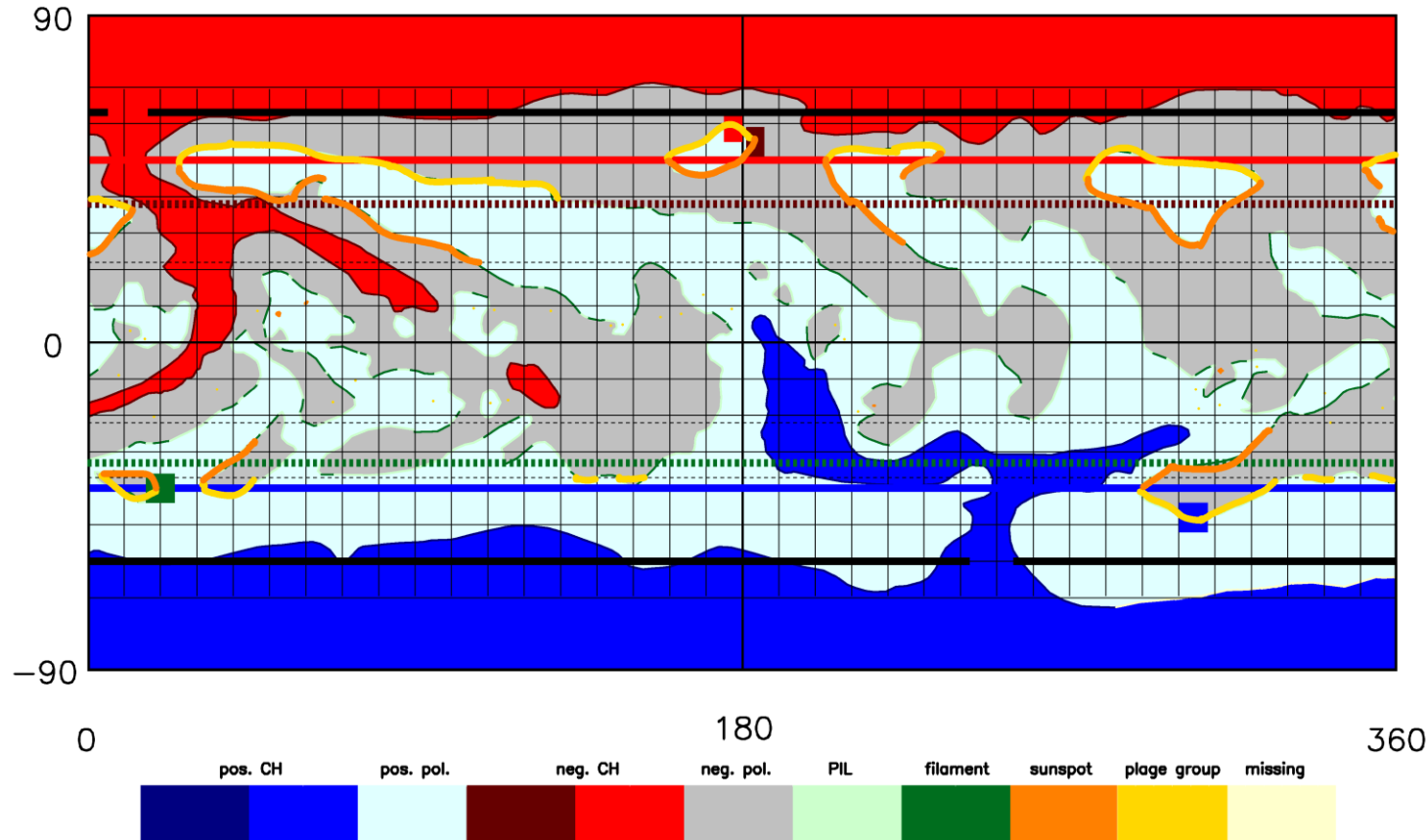
End date (longitude=0):1984-08-13T19:52:39

Start date (longitude=360):1984-07-17T14:36:16

B angle end date 6.600

B angle start date 4.620

CR1751



Solar minimum Jul-Aug 1984

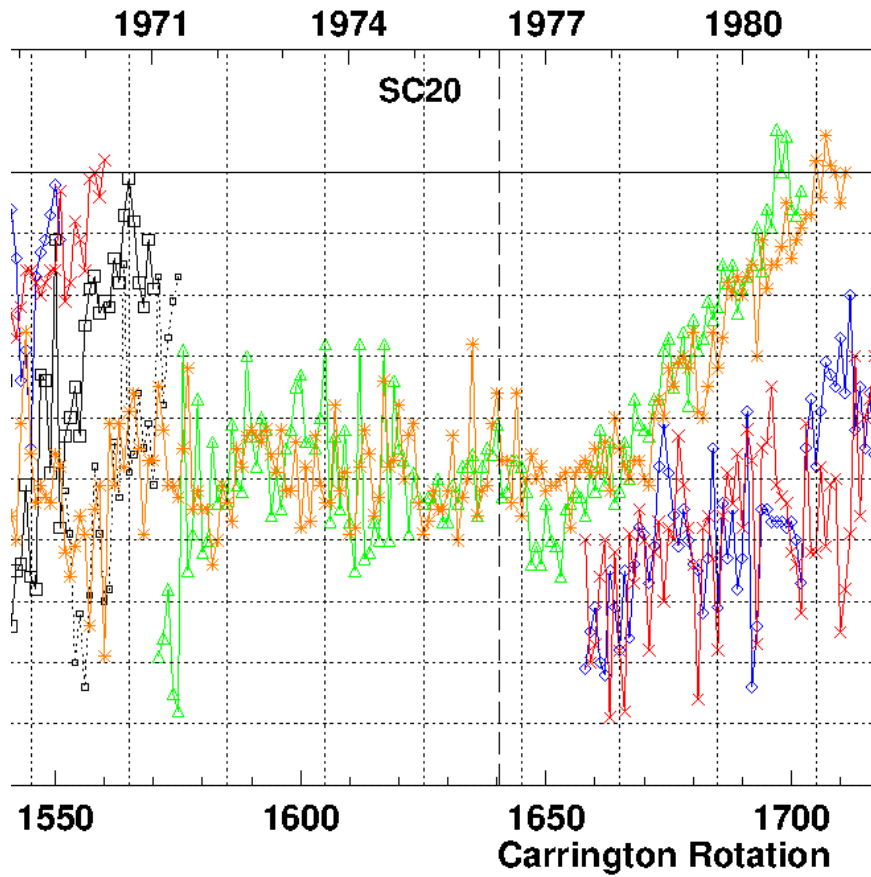
Polar coronal hole (CH)  
latitudinal extent  $\sim 28-30^\circ$ ,  
with low-latitude extensions

Primary PIL (polarity  
inversion line)  $50.0^\circ\text{N}$ ,  $40.5^\circ\text{S}$

Secondary PIL (polarity  
inversion line)  $38.5^\circ\text{N}$ ,  $33.5^\circ\text{S}$

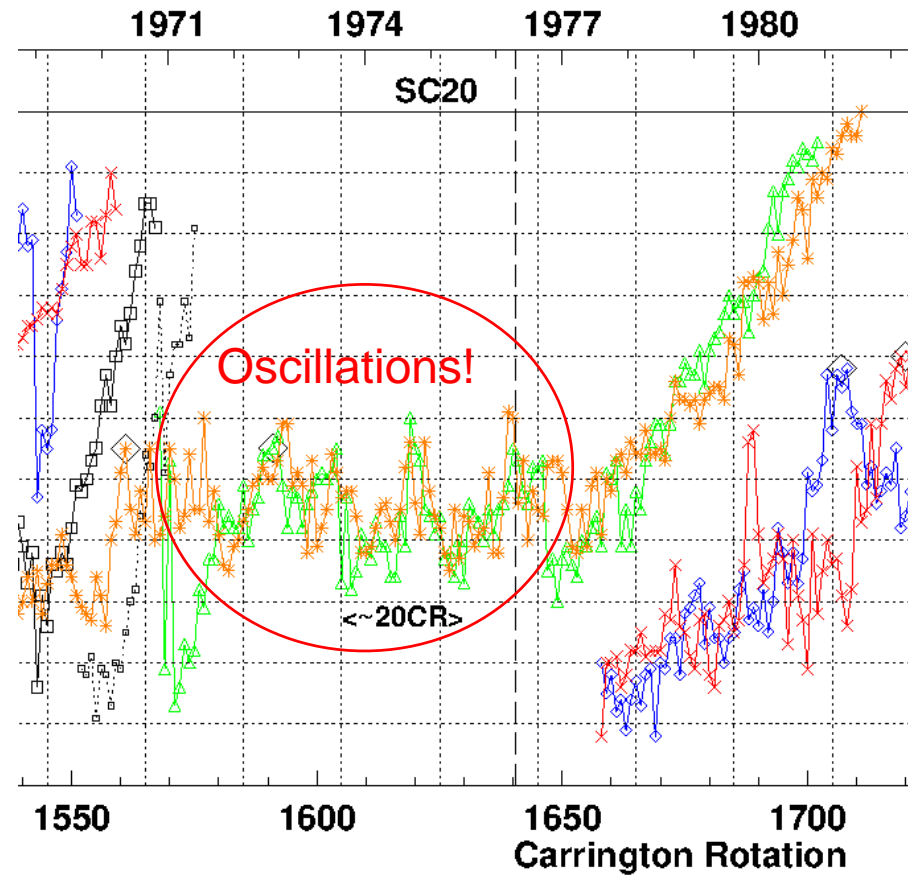
# Median PILs (Polarity Inversion Lines) Cleaner, Better Story than Filaments

## Maximum Polar Crown Filaments



$\pm 55^\circ$

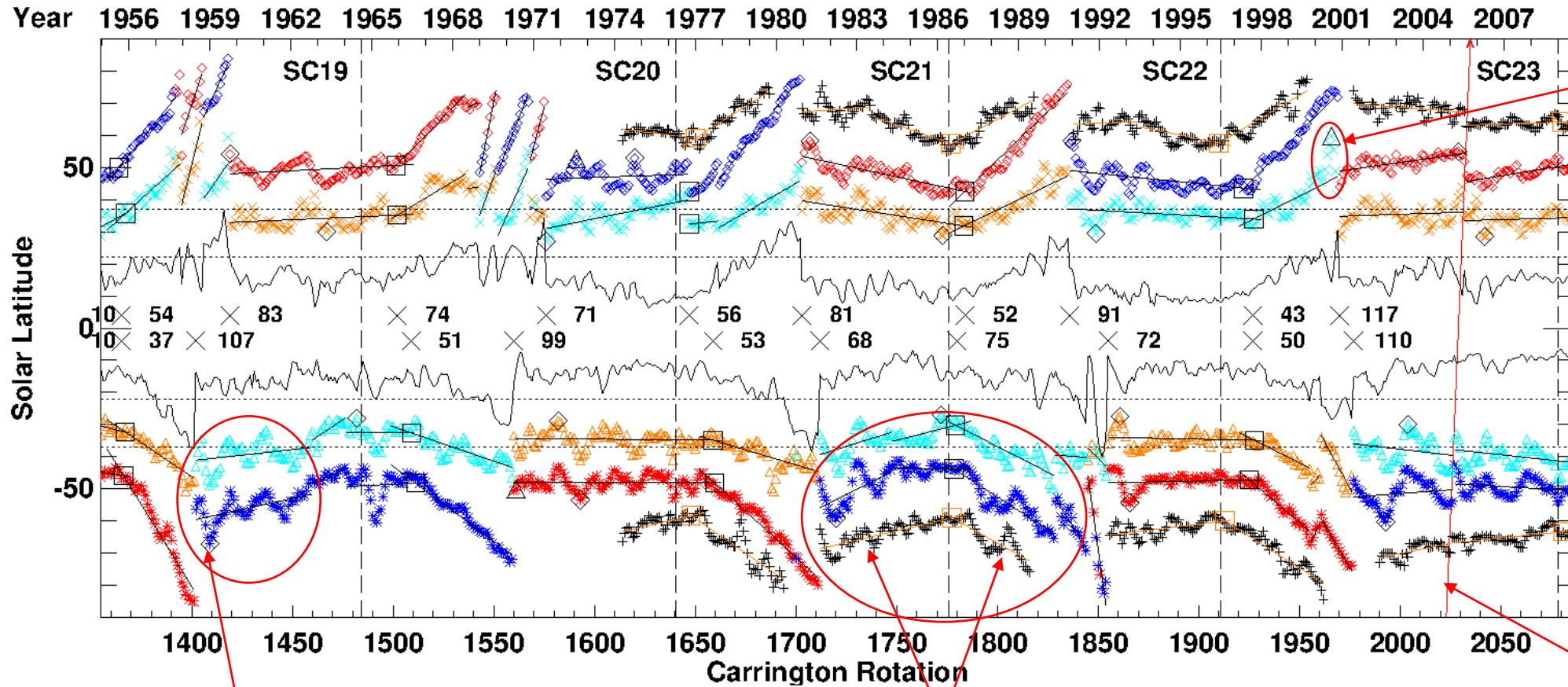
## Median Polar Crown PILs



# SC 19-23 Oscillations after Secondary PIL Rush-to-the-Pole End Peaks

The end of the Rush-to-the-Pole (RttP) of the secondary PIL (diamonds or triangles) is usually later than the RttP end of the primary PIL marking the polarity change at the pole at Solar maximum (Smax), where red=negative and blue=positive

**Median Smoothed Polar CH Boundaries and PILs and Differences Between Primary and Secondary PILs**



Early end  $\sim 60^\circ\text{N}$  in the Rush-to-the-Pole (RttP) Secondary PIL

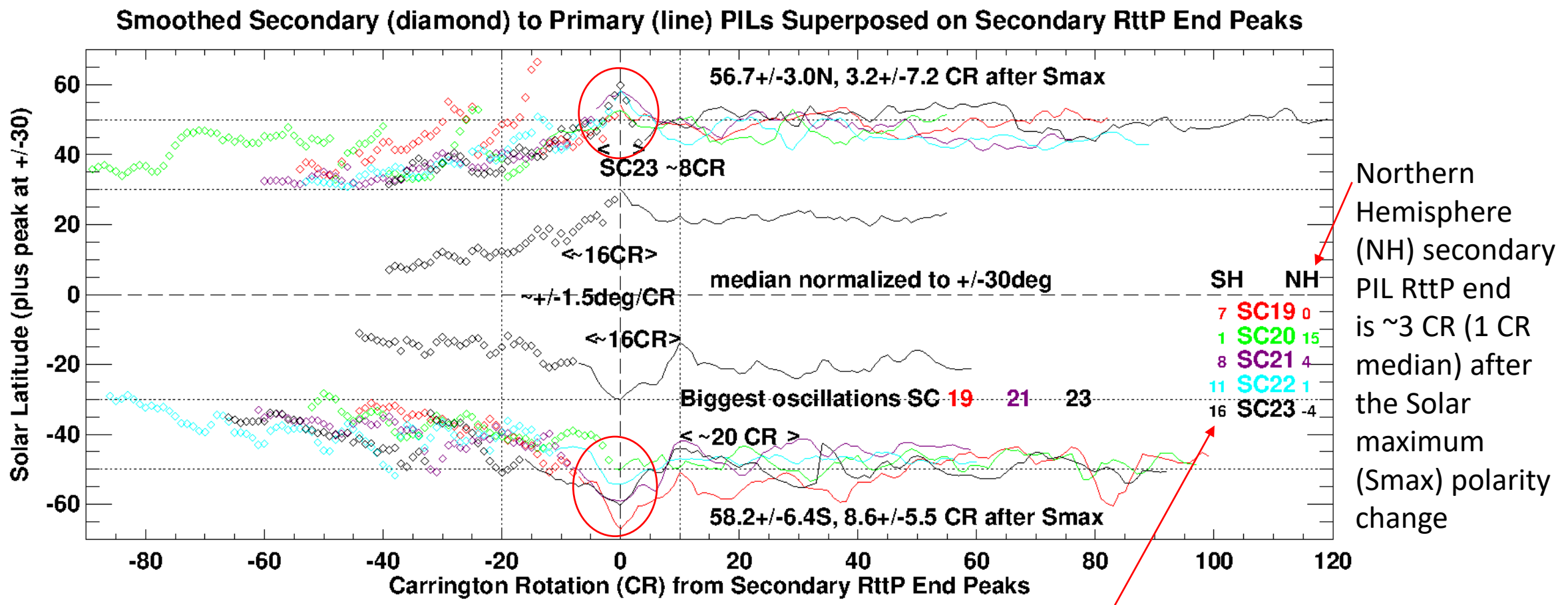
Smax Secondary Rush-to-the Pole (RttP) PIL ends in Primary non-Rush PIL  $\sim 67^\circ\text{S}$

Oscillations also in CH boundary

Equatorward drop in latitude from South to North  $\sim 15^\circ/\text{CR}$  (CR=Carrington Rotation)

# Secondary PIL Ends Its Rush-to-the-Pole (RttP) $\sim 57^\circ\text{N}$ ( $\sim 58^\circ\text{S}$ ) in the Primary non-Rush PIL $\sim 3$ CR ( $\sim 9$ CR) after Solar Maximum Polarity Change

Superposed Epoch Plot with the end of the Secondary PIL's RttP at Zero

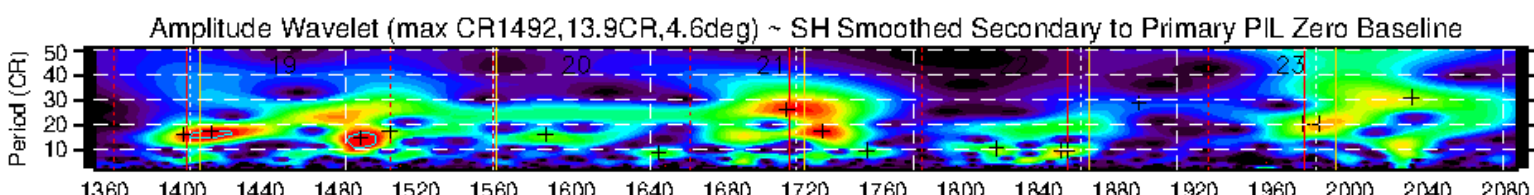
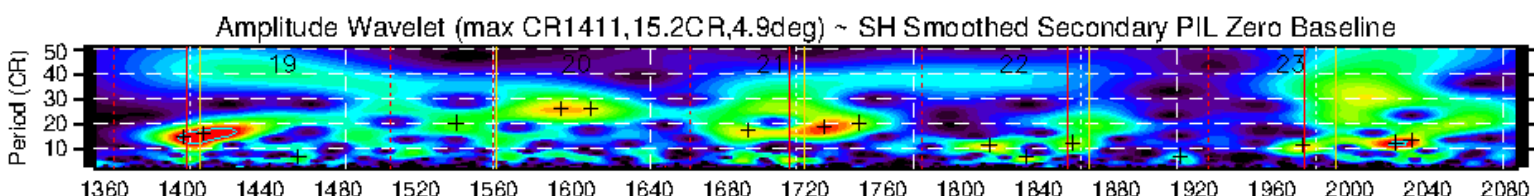
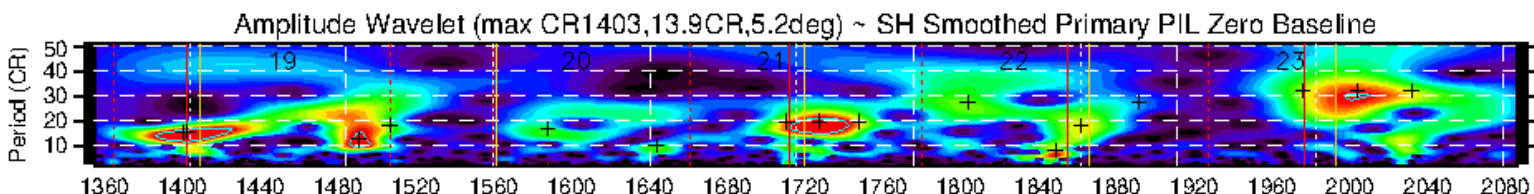
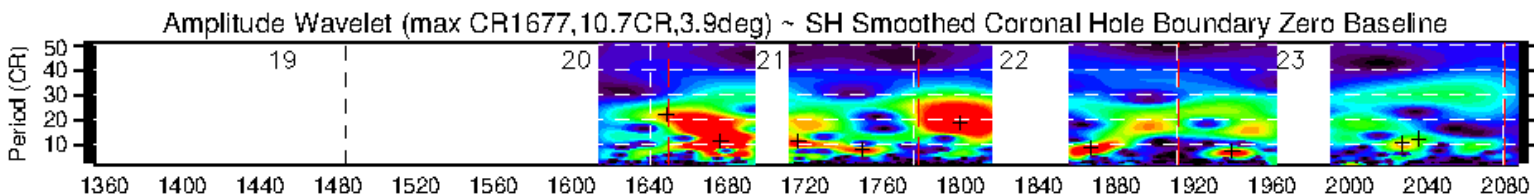
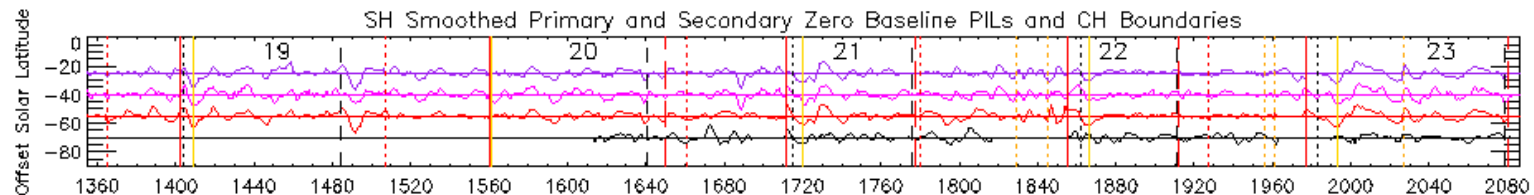


Southern Hemisphere (SH) secondary PIL RttP end is  $\sim 9$  CR (8 CR median) after the solar maximum polarity change

# ~16 CR Oscillations from Solar Maximum Transition Peaks

Red lines Smax  
polarity change.

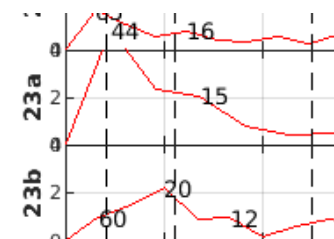
Gold lines  
transition peaks.



~10-35 CR  
period of  
oscillations

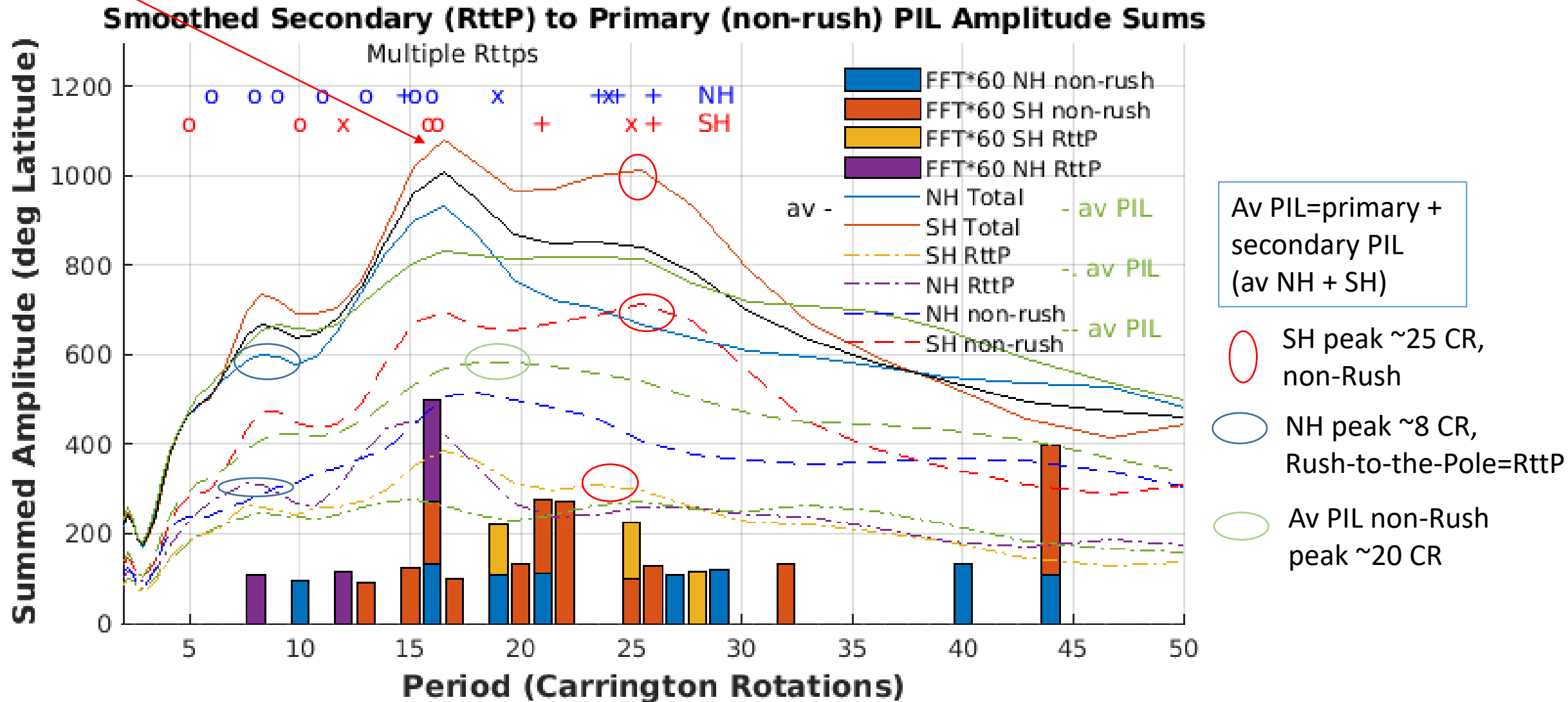
Largest SH (Southern Hemisphere) oscillations in odd SCs 19, 21, 23

Fast Fourier Transform  
(FFT) Amplitudes  
(°Lat) and Periods (CR)  
SC 23a ~5° ~35 CR  
SC 23b ~2° ~20-25 CR



Oscillations ~16 CR at  
Smax Transition Peaks

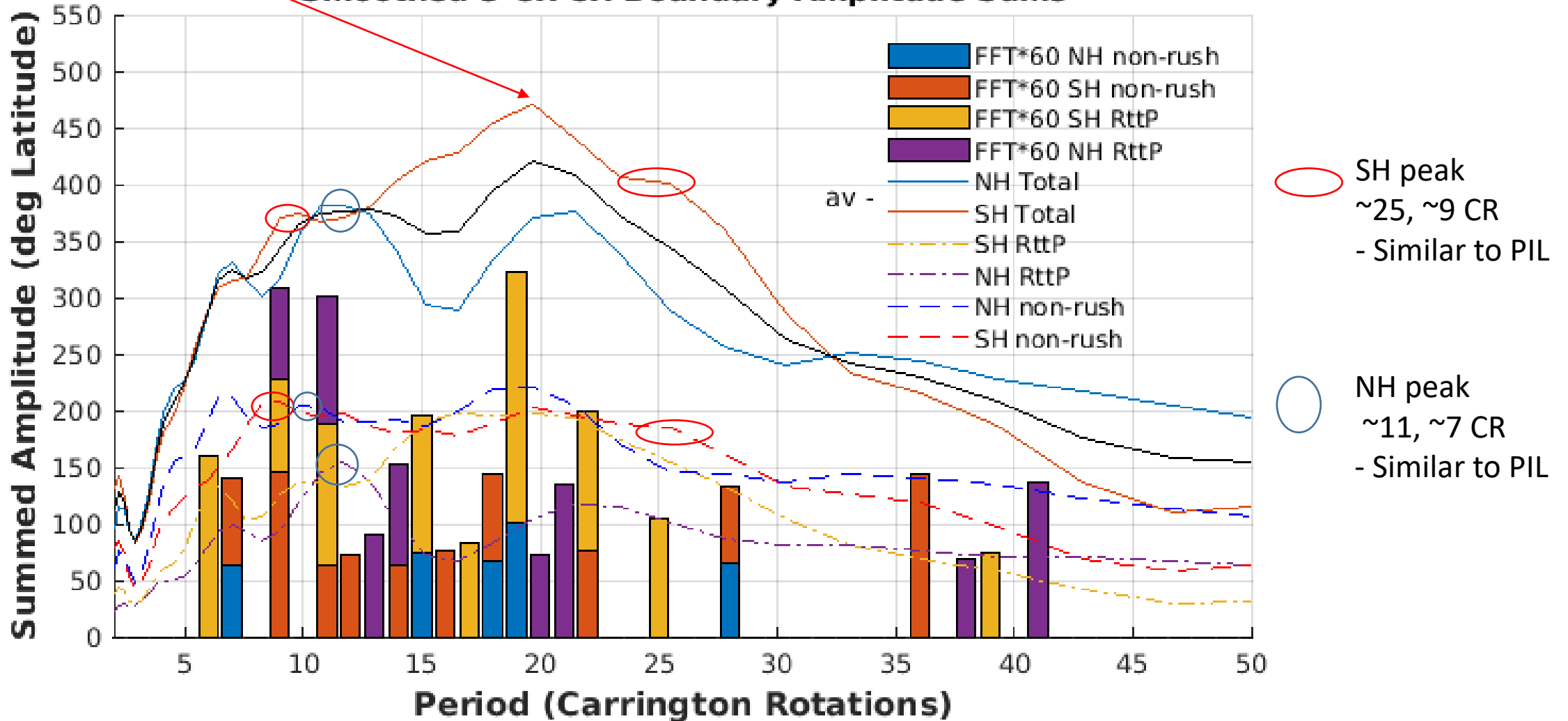
# ~16 CR PIL Peak Period from FFT Bars and Wavelet Amplitude Sums



~20, ~10 CR Periods for CH Boundary from FFT Bars and Wavelet Sums

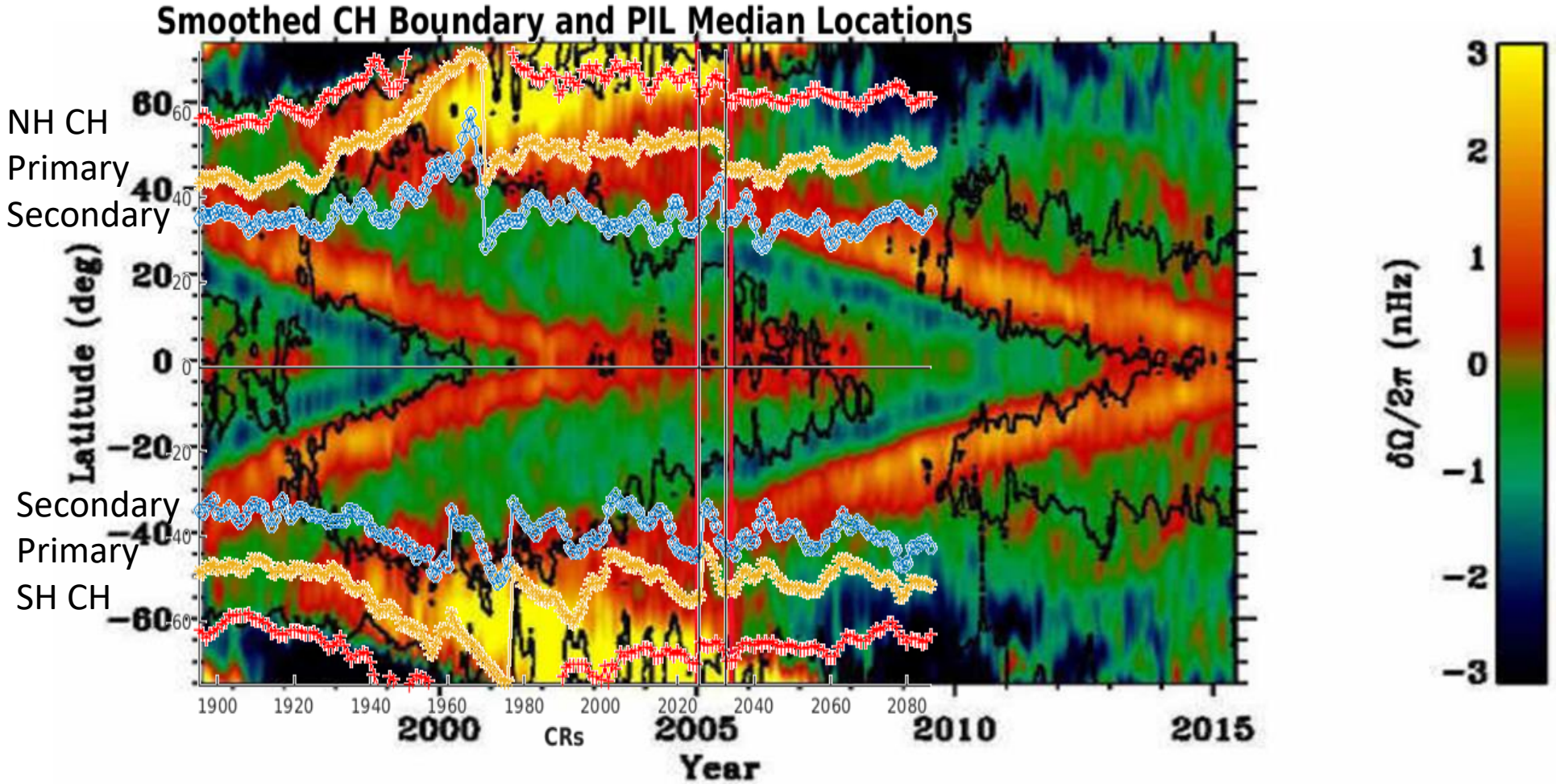
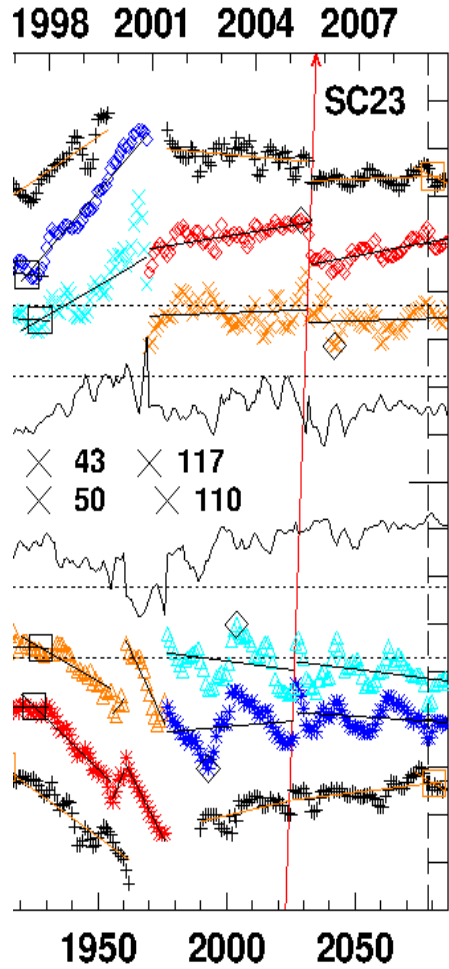
No major ~16 CR period because have only one CH Boundary

Smoothed 3-CR CH Boundary Amplitude Sums





# ~15°/CR Disturbance Expands Polar Coronal Holes ~50% from South to North in 2005 CR2023-35



Southern PIL wavelets ~30 CR before, ~22 CR after and smaller amplitudes

From Howe (2016) meridional flows and 5 G unsigned magnetic flux. Overplot SC 23 PILs and Coronal Hole boundaries.

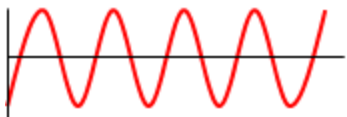
~±57° Latitude



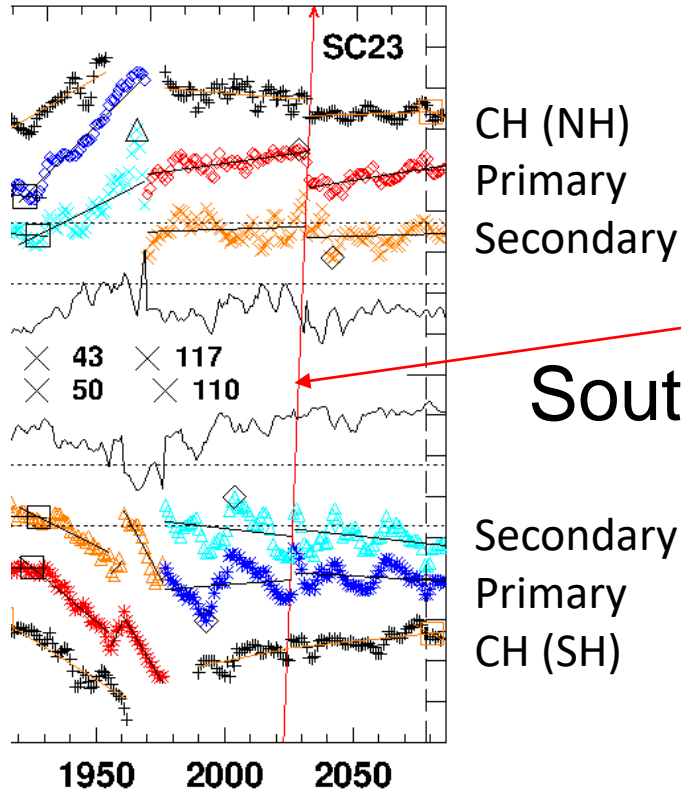
# CONCLUSIONS

The RttP End of the Secondary PIL in the Transition to the non-rush Primary PIL



  
~16 CR Oscillations in Solar Latitude!

1998 2001 2004 2007



*Other periods shared with Coronal Hole (CH) boundaries of ~25 (SH), ~20 (non-rush), ~7-11 CR (mostly NH)*

**AWESOME** ~15°/CR Disturbance from Southern Hemisphere (SH) to Northern Hemisphere (NH) in 2005 SC 23 solar minimum!

*What do modelers think happened here?*

<https://www2.hao.ucar.edu/mcintosh-archive/four-cycles-solar-synoptic-maps>

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